Nuclear Physics Highlights from Stable Ion Beam Facilities ECOS -2011-2014



## **Plan of the talk**



A variety of Physics Q's (suggested by the labs)

Tools and phenomena Selectivity and sensitivity

Tomorrow ?





P.T. Greenlees<sup>1,\*</sup> I. Rubert<sup>2</sup> I. Piot<sup>2</sup> B. I.P. Gall<sup>2</sup> I. L. Andersson<sup>3</sup> M. Asai<sup>4</sup> Z. Asfari<sup>2</sup> D. M. Cox<sup>3</sup> F. Dechery



- Solid angle acceptance (central m/q and energy) 10 msr
- Typical transmission ~12% per charge state

## de la

### Fission life-times : a link towards the production of SHE

![](_page_5_Picture_2.jpeg)

![](_page_5_Figure_3.jpeg)

M.O. Frégeau et al., Phys. Rev. Lett. 108, 12270 (2012)

• $X_K$  fluorescence of the atom with 120 protons has been observed.

 Compound nucleus lifetime of the order of the K-vacancy lifetime (~10<sup>-18</sup>s)

### New Collaboration: GANIL-IPNO-SPhN?-ANU

*Measure in coincidence with X-rays The angular distribution of fragments* 

![](_page_5_Figure_9.jpeg)

•Determination of the fusion components on  $4\pi$  from X-ray fluorescence Reconcile the most probable fission times inferred from angular distributions with the average fission times inferred from direct measurements

![](_page_6_Figure_0.jpeg)

200

Ex (MeV)

E (MeV)

Applications reactor heat, radio toxicity control

# Isotopic distribution of fission fragments : OpenF. Farget et al., GANILQuestions

![](_page_7_Figure_1.jpeg)

What happens when an exotic nuclei undergoes fast rotation Is this the next question we want to address

![](_page_8_Figure_1.jpeg)

![](_page_8_Figure_2.jpeg)

## **Prompt spectroscopy of M and Z identified fission frag**

![](_page_9_Figure_1.jpeg)

<sup>238</sup>U+<sup>9</sup>Be at 6 Mev/U Tomorrow AGATA@GANIL

> Thus uniquely identify M and ZDoppler correction for the emitted  $\gamma$  rays  $\vec{v}$  of the fragment & angle of the segment of the clover detector

A. Navin and M. Rejmund McGraw-Hill Yearbook of Science & Technology (2014) pg

![](_page_9_Figure_5.jpeg)

![](_page_10_Figure_0.jpeg)

![](_page_10_Figure_1.jpeg)

![](_page_11_Figure_0.jpeg)

![](_page_11_Figure_1.jpeg)

200

100

Coun

150

100

\_\_\_\_\_

100

Neutron Rich <sub>45</sub>Rh

Work in progress

Lithium Inverse Cinematiques ORsay Neutron source

![](_page_12_Figure_1.jpeg)

### Isoscalar Component of the PDR for Nuclei with n-excess

![](_page_13_Figure_1.jpeg)

Inelastic scattering of <sup>17</sup>O @ 20 MeV/u on different targets + y-rays in coincidence

![](_page_13_Picture_3.jpeg)

### TWO EXPERIMENTS PERFORMED:

TRACE

TKE [MeV]

Studied Nuclei: 208Pb 90Zr
 R. Nicolini (Università di Milano /INFN)
 D.Mengoni (Università di Padova/INFN)

Studied nuclei: 208Pb, 124Sn,140Ce
 M. Kmiecik (IFJ PAN Kraków) ,
 F. Crespi (Univ. di Milano/INFN)

F.C.L. Crespi et al., EPJ WoC, INPC 2013

### Pygmy Dipole Resonance in <sup>124</sup>Sn by inelastic scattering of <sup>17</sup>O L. Pellegri Phys. Lett. B (in press)

Dominant Isoscalar excitation: n and p transition densities are **in phase** inside the nucleus, at **the surface only the n-part survives** Inelastic scattering: Interaction Surface Peaked (<sup>17</sup>O)

![](_page_14_Figure_2.jpeg)

The emitted  $\gamma$  rays were detected with high resolution with the AGATA demonstrator array and the scattered ions were detected in two segmented  $\Delta E - E$  silicon telescopes.

## Isospin character of low-lying pygmy states in <sup>208</sup>Pb via inelastic scattering of <sup>17</sup>O F.C.L. Crespi et al., PRL 113, 012501 (2014)

![](_page_15_Figure_1.jpeg)

Isospin Character of Low-Lying Pygmy Dipole States in <sup>208</sup>Pb via Inelastic Scattering of <sup>17</sup>O Ions

The *E*1 transitions cross sections for <sup>208</sup>Pb were analyzed for the first time using a microscopic form factor and the isoscalar potential was found to depend on the presence of the neutron skin. A

![](_page_16_Picture_5.jpeg)

Pygmy dipole resonance in  $^{124}\mathrm{Sn}$  populated by inelastic scattering of  $^{17}\mathrm{O}$ 

For the 1<sup>-</sup> states a DWBA analysis based on a microscopically calculated form factor was performed and showed a sensitivity to the surface part of the transition density. Being the transition density dominated on the surface by the neutron component one can deduce that the pygmy states <sup>124</sup>Sn are associated with the excitation of surface neutrons, mainly those in the neutron skin. Therefore in the future it will be very interesting to perform these studies on the isotopic chain of Sn and other neutron-rich nuclei also using other probes as protons at medium energy. Cyclotron Center Bronowice Institute of Nuclear Physics Polish Academy of Sciences Krakow, Poland

## A) Study of collective modes excited by high-energy protons

![](_page_17_Picture_2.jpeg)

 HECTOR array to measure high energy gammarays

2. **KRATTA** array (triple Csl telescopes) at forward direction to measure the energy of inelastically scatter protons

3\*. **PARIS** Demonstrator for PDR measurements with high-resolution, in coincidence with KRATTA

![](_page_17_Picture_6.jpeg)

![](_page_17_Picture_7.jpeg)

![](_page_18_Picture_0.jpeg)

#### Hot GDR study in nuclei in the mass region A~120-132 with MEDEA at LNS

Onset of the quenching of the Giant Dipole Resonances at high excitation energies

E\* A<sub>res</sub> <sup>116</sup>Sn+<sup>12</sup>C 17A MeV 150 124 190 123 <sup>116</sup>Sn+<sup>12</sup>C 23A MeV 132 270 <sup>116</sup>Sn+<sup>24</sup>Mg 17A MeV

D.Santonocito<sup>a</sup>), Y.Blumenfeld<sup>b</sup>), C. Agodi<sup>a</sup>), R.Alba<sup>a</sup>), G. Bellia<sup>a</sup>),<sup>c</sup>), R. Coniglione<sup>a</sup>), F. Delaunay<sup>b)</sup>\*, A. Del Zoppo<sup>a)</sup>, P. Finocchiaro<sup>a)</sup>, F. Hongmei<sup>a)</sup>, V. Lima<sup>b)</sup>, C. Maiolino<sup>a)</sup>, E. Migneco<sup>a),c)</sup>, P.Piattelli<sup>a)</sup>, P. Sapienza<sup>a)</sup>, J.A. Scarpaci<sup>b)<sup>†</sup></sup>, O.Wieland<sup>d)</sup>

The evolution of the Giant Dipole Resonance properties in nuclei of mass A =120-132 has been investigated in an excitation energy range between 150 and 270 MeV through the study of complete and nearly complete fusion reactions. Evidence of a quenching of the GDR gamma yield was found at 270 MeV excitation energy.

A limiting excitation energy for the collective motion of about  $E^*/A \sim 2$ **MeV/A** was exctracted.

![](_page_18_Figure_7.jpeg)

ao mass  $A \sim 132$ . Evidence of a limiting excitation for the collective motion was also extracted in nuclei of mass  $A = 60 \div 70$  but the value of about  $E^*/A \simeq 5$  MeV, differs significantly from the ones measured for nuclei in the mass region  $A = 105 \div 135$  suggesting the existence a mass dependence of the limiting excitation energy per nucleon, for the collective motion. Interesting similarities in trend and absolute values can be found when comparing the limiting excitation energy for the collective motion with the energy at which the plateau of the caloric curve sets in, indicating the onset of liquid-gas phase transition. This feature suggests a possible link between GDR disappearance and liquid-gas phase transition which deserves further investigation from both theoretical and experimental points of view. In particular extending the systematics of the GDR to hot nuclei with  $A = 160 \div 180$  could provide further information on the 5 possible link between GDR disappearance and liquid-gas phase transition and therefore shed additional light on the mechanism responsible for the GDR quenching. a) Solid symbols re

![](_page_18_Figure_9.jpeg)

represent the spectra after bremsstrahlung subtraction.

b) Statistical gamma spectra compared to CASCADE calculations shown as red lines.

c) Linearized Spectra compared to Lorentzian function used in the calculation for each reactions.

## <sup>174</sup>W: ORDER-TO-CHAOS TRANSITION

### QUASI-CONTINUUM $\gamma$ - $\gamma$ MATRICES

High-Spin Fusion Evaporation  ${}^{50}$ Ti on  ${}^{128}$ Te @ 217 MeV, I  $\geq$  60 $\hbar$ 

![](_page_19_Figure_4.jpeg)

Goal: populate <sup>174</sup>W at the **highest possible spins** ( $\geq$ 60ħ), in order to make the **statistical fluctuation analysis of the ridge-valley structures in the**  $\gamma$ - $\gamma$  **matrices**, to estimate the number of low-*K* and high-*K* bands and their correlation

PHYSICAL REVIEW C 88, 034312 (2013)

HIGH TEMP **4 Triple Clusters** 2 and 3 folds:  $\varepsilon_{2\gamma}=30\%, \varepsilon_{3\gamma}=10\%$ ( $M_{\gamma}=30$ )

> region. The present results suggest that a K-mixing process due to temperature effects plays an important role already at rather low excitation energy, namely in the onset region of band mixing, here probed by a global analysis of decay properties of the entire body of discrete excited bands. This represents a step forward in the understanding of the basic

![](_page_19_Figure_9.jpeg)

V. Vandone et al. Global properties of K hindrance probed by the  $\gamma$  decay of the warm rotating <sup>174</sup>W nucleus

## SHAPE TRANSITION IN THE OS ISOTOPES

![](_page_20_Figure_2.jpeg)

### Shape evolution in the n-rich Os isotopes: prompt γ-ray spectroscopy of <sup>196</sup>Os P.R. John et al. PRC 90 021301 (2014) rapid comm.

![](_page_21_Figure_1.jpeg)

### Superdeformation in A~40 nuclei

E. Ideguchi (RCNP, Osaka Univ.) – January 2013

Search for superdeformed bands in  $^{35,36}$ S,  $^{40}$ Ar via  $^{18}$ O +  $^{26}$ Mg ->  $^{44}$ Ca\*

![](_page_22_Figure_3.jpeg)

Fig. 1.  $\gamma$ -ray energy spectrum gated 13 02-keV transition. The 1576-, 1302- and 978-keV transitions belong to the same band transition. The peaks 1 beled with the triangles are new transitions of <sup>35</sup>S.

![](_page_22_Figure_5.jpeg)

**Fig. 2.** Comparison between experimental and simulated spectra for the residual Doppler shifted 1576-keV peak in the forward and backward angles of the germanium detectors.

Investigations of the orbitals responsible for superdeformation in the mass region, f7/2, nearby are important to infer the existence of superdeformed states in  $^{32}$ S.

The de-excitation of the superdeformed band was observed from  $19/2^-$  to  $7/2^-$ . The superdeformed band structure was highlighted by the measurement of half-life and relative intensities of the intra-band transition. The  $f_{7/2}$  negative parity intruder orbital seems to be responsible for the negative

#### Shintaro GO (RCNP, Osaka University), ARIS2014 Proceedings

### Time Dependent Recoil In Vacuum on H-like ions: <sup>24</sup>Mg - revisited

G. Georgiev (CSNSM), A.E. Stuchbery (ANU), Dec. 2012

10

12

16

14 N=Z 18

20

*Experiment:* High accuracy (< 2%) model independent (*B from first principles*) g-factor value for short-lived (ps) excited states Theory: g-factors of the 2<sup>+</sup> states in N=Z nuclei should be slightly higher than 0.5 8-fold segmented 0.70 annular detector 0.65 **OUPS** <sup>24</sup>Mg+<sup>93</sup>Nb coulex (Orsay Universal Plunger System) (<u>+</u>) 0.60 0.6 0.55 g(2<sup>+</sup>) 0.5 0.50 20 60 80 100 120 40 0 flight time (ps) A. Kusoglu, A. Stuchbery, G. Georgiev et al. 0.4

*Our result:* first experimental evidence of deviation from g=0.5  $\rightarrow$  stringent test of the nuclear theories

## Nuclear spin orientation in incomplete fusion reactions

Nuclear spin orientation – a must for nuclear moments studies

- Fusion-evaporation reactions 25 % 75 % alignment
- Projectile-fragmentation 8 % 13 %
- Direct reactions (single-nucleon transfer) ~ 13 %
- Incomplete fusion (multi-nucleon transfer?) ???

![](_page_24_Figure_6.jpeg)

Amplitude = 8 (1) % Spin alignment = 23 (3) %

Amplitude = 4.8 (8) % **Spin alignment = 12.5 (20) %** 

**Results:** 

- considerable spin alignment in <sup>7</sup>Li induced reactions;
- dependence on the number of transferred nucleons?

G. Georgiev (CSNSM) Dec. 2013 <sup>68</sup>Ga <sup>69</sup>Ga <sup>70</sup>Ga <sup>71</sup>Ga <sup>66</sup>Zn <sup>67</sup>Zn <sup>68</sup>Zn <sup>69</sup>Zn <sup>70</sup>Zn <sup>65</sup>Cu <sup>66</sup>Cu <sup>67</sup>Cu <sup>68</sup>Cu <sup>69</sup>Cu <sup>64</sup>Ni <sup>65</sup>Ni <sup>66</sup>Ni <sup>67</sup>Ni <sup>68</sup>Ni

![](_page_24_Figure_13.jpeg)

<sup>7</sup>Li+<sup>64</sup>Ni Triton 1n , triton1pn

## Spontaneous chiral symmetry breaking in <sup>124</sup>Cs

The central European Array for Gamma Levels Evaluation

- Experiment EAGLE [1] array.
- DSA method measurement leading to B(M1) and B(E2) determination.
- Spontaneous chiral symmetry breaking [2] is proven if one can see two rotational bands (chiral partner bands) with similar reduced transition probabilities, (B(M1) and B(E2)).
- Preliminary results for <sup>124</sup>Cs - B(M1) in partner bands.

[1] J. Mierzejewski et al., NIM A 659, 84 (2011).[2] E. Grodner et al., Eur. Phys. J A27, 325 (2006).

![](_page_25_Picture_7.jpeg)

![](_page_25_Figure_8.jpeg)

### Neutron pair transfer in <sup>60</sup>Ni+<sup>116</sup>Sn far below the Coulomb barrier

![](_page_26_Figure_1.jpeg)

## Transfer strength very close to the g.s. to g.s. transitions

![](_page_26_Figure_3.jpeg)

The experimental transfer probabilities are well reproduced, for the first time with heavy ion reactions, in absolute values and in slope by microscopic calculations which incorporate nucleon-nucleon pairing correlations

D.Montanari et al., PRL113(2014)052501

![](_page_27_Picture_0.jpeg)

## **Pairing correlations with MAGNEX**

<sup>18</sup>O+<sup>13</sup>C at 84 MeV

![](_page_27_Figure_2.jpeg)

## Barrier distributions for fusion

ŚRODOWISKOWE LABORATORIUM CIĘŻKICH JONÓW

![](_page_28_Figure_2.jpeg)

Recent papers: E. Piasecki et al. PRC 85 (2012) 054604 and 054608

## NAP

### Stellar burning rates and <sup>14</sup>N(p,g)<sup>15</sup>O reaction

![](_page_30_Figure_1.jpeg)

slowest reaction => it
 determines the rate of
 the CNO cycle
 → age of globular clusters
 → Solar Composition
 Problem

The sub-threshold resonance corresponding to the <u>first excited 3/2<sup>±</sup></u> <u>state in <sup>15</sup>O</u> is predicted to play a dominant role when extrapolating the cross-section to the Gamow peak region

### width of the resonance <=> lifetime of the nuclear state

First application of the high gamma energy resolution and position sensitivity of AGATA to investigate ≈fs nuclear level lifetimes

### Lifetime measurement of the 6.79 MeV state in <sup>15</sup>O with the AGATA Demonstrator

![](_page_30_Figure_7.jpeg)

4 asymmetric triple-clusters 12 36-fold **segmented** HPGe

first interaction point and γ energy event-by-event: Lineshape analysis with a few degrees resolution

### Energy vs theta first interaction point

![](_page_31_Figure_1.jpeg)

### Lineshape analysis for the 6.79 MeV state in <sup>15</sup>O

![](_page_31_Figure_3.jpeg)

C. Michelangelo (pvt Comm)

theta [deg]

#### Nuclear astrophysics @ LNS – Trojan Horse Method Applications

![](_page_32_Figure_1.jpeg)

Red line Rmatrix based on THM data Points, direct measurements

First time in Gamow window (<500 keV)

Reaction rate is calculated in the AGB star temperature window for the first time. The ratio to the NACRE evaluation is reported in panel b

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THE ASTROPHYSICAL JOURNAL LETTERS, 739:L54 (6pp), 2011 October 1 © 2011. The American Astronomical Society. All rights reserved. Printed in the U.S.A.
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A. Tumino et al . Acta Phys.Pol. B45, 181 (2014)

THM is an indirect method that allow to extract nuclear information of a two-body reaction, even at astrophysical energies, by means of the quasi-free contribution in an appropriate three-body reaction.

 $A + x \rightarrow C + c$ ;  $A + a \rightarrow C + c + b$ ;

<sup>19</sup>F(p,α)<sup>16</sup>O (two-body reaction) studied at LNS via <sup>19</sup>F(d,αn)<sup>16</sup>O (three-body reaction) At 50 MeV @Tandem

![](_page_32_Figure_10.jpeg)

Search for new resonant states in <sup>10</sup>C and <sup>11</sup>C and their impact on the cosmological lithium problem

F. Hammache, N. de Sereville, I. Stefan

Primordial nucleosynthesis (BBN) is one of the three evidences for the Big-Bang model

When T <  $10^9 \text{ K} \rightarrow \text{BBN starts}$ 

- Production of D, <sup>3</sup>He, <sup>4</sup>He, <sup>7</sup>Li
- Abundances depend on baryonic density
- D, <sup>3</sup>He, <sup>4</sup>He, observations agree with predictions (BBN + CMB) Metal poor halo dwarf stars

<sup>7</sup>Li problem:  $(^{7}Li/H)_{BBN} / (^{7}Li/H)_{obs} = 4$ 

### Possible explanations:

- Physics beyond standard model: super-symmetry, constant variation, ....
- Observations: can <sup>7</sup>Li be uniformly destroyed in the Splite plateau region?
- Nuclear physics: <sup>7</sup>Li produced by <sup>7</sup>Be EC & <sup>3</sup>He(<sup>4</sup>He,  $\gamma_{o}$ )<sup>7</sup>Be known better than 15%

### Last proposed solution studied with SPLITPOLE @ IPN Orsay

Finally, our two results concerning <sup>10</sup>C and <sup>11</sup>C compound nuclei put an end to the various discussions concerning the missing resonant states in these nuclei, which were thought to partially or totally solve the <sup>7</sup>Li problem [19–21] and exclude <sup>7</sup>Be + <sup>3</sup>He and <sup>7</sup>Be + <sup>4</sup>He reaction channels as responsible for the observed <sup>7</sup>Li deficit. Direct kinematics <sup>70</sup>Zn(d,<sup>3</sup>He)<sup>69</sup>Cu with 27 MeV deuteron beam and SPLITPOLE spectrometer

Part of the program: Systematic of the evolution for  $f_{7/2}$  proton-hole strength from stability to <sup>78</sup>Ni ?

![](_page_34_Figure_2.jpeg)

![](_page_35_Figure_0.jpeg)

Spectroscopy of proton-rich  ${}^{66}$ Se up to  $J^{\pi} = 6^+$ : Isospin-breaking effect in the A = 66 isobaric triplet

Timescale of the multifragmentation is directly reflected in the shape of the largest fragment charge/mass distribution

#### INDRA collaboration +M. Ploszajczak (GANIL) +R. Botet (LPS Orsay)

D. Gruyer et al., Phys. Rev. Lett. 110, 172701 (2013)

![](_page_36_Figure_3.jpeg)

### Studies of heavy ions reactions with the CHIMERA@LNS

#### CHIMERA@LNS

![](_page_37_Picture_2.jpeg)

 Reaction Dynamics at Fermi energy (\*)
 EOS - density dependence of the symmetry term (\*\*)
 Reactions and Structure with Radioactive Fragmentation Beam (\*\*\*)
 Correlation and interferometry : FARCOS (\*\*\*\*)
 DSSD(300+1500+Csl)

\*\*\*) In-Flight RIB production

### (\*) Dyn/Stat Fission: Size or Isospin effects?

![](_page_37_Figure_6.jpeg)

Stochastic Mean Field (SMF) +GEMINI  $E_{sym} \approx (\rho / \rho_0)^2$   $I_4 \qquad \qquad Primary fragments$   $I_5 \qquad \qquad Primary fragments$   $I_7 \qquad \qquad Primary fragmen$ 

(\*\*)Comparison of IMF data with

E.De Filippo et al, PRC86 014610 2012

![](_page_37_Figure_9.jpeg)

(\*\*\*\*) CHIMERA + FARCOS correlator

![](_page_37_Picture_11.jpeg)

![](_page_38_Picture_0.jpeg)

![](_page_38_Figure_1.jpeg)

## B) Study of the dynamics of few-body systems

1. Detector BINA (moved from KVI Groningen) for light nuclei reactions studies

Wall:

- MWPC (3 planes)
- ∆E (24 x 2 mm)
- E (20 x 120 mm)
  Ball:
- Phoswich (149 x 90/30 mm)

![](_page_39_Picture_7.jpeg)

![](_page_39_Picture_8.jpeg)

# CCB Krakow and HIL Warsaw are in the ENSAR2 H2020 project as TNA facility

### ECOS: a Path to ...

What can we different that we cannot do today?

What are the physics questions that we specifically that we want to address for which we need very intense stable beams ?

What are the things we need to think and develop (other than the machine)

![](_page_40_Picture_4.jpeg)

To Boldly Go Where No Man

(Other speaker are going to answer this)

![](_page_40_Picture_6.jpeg)

**Evolution to revolution**